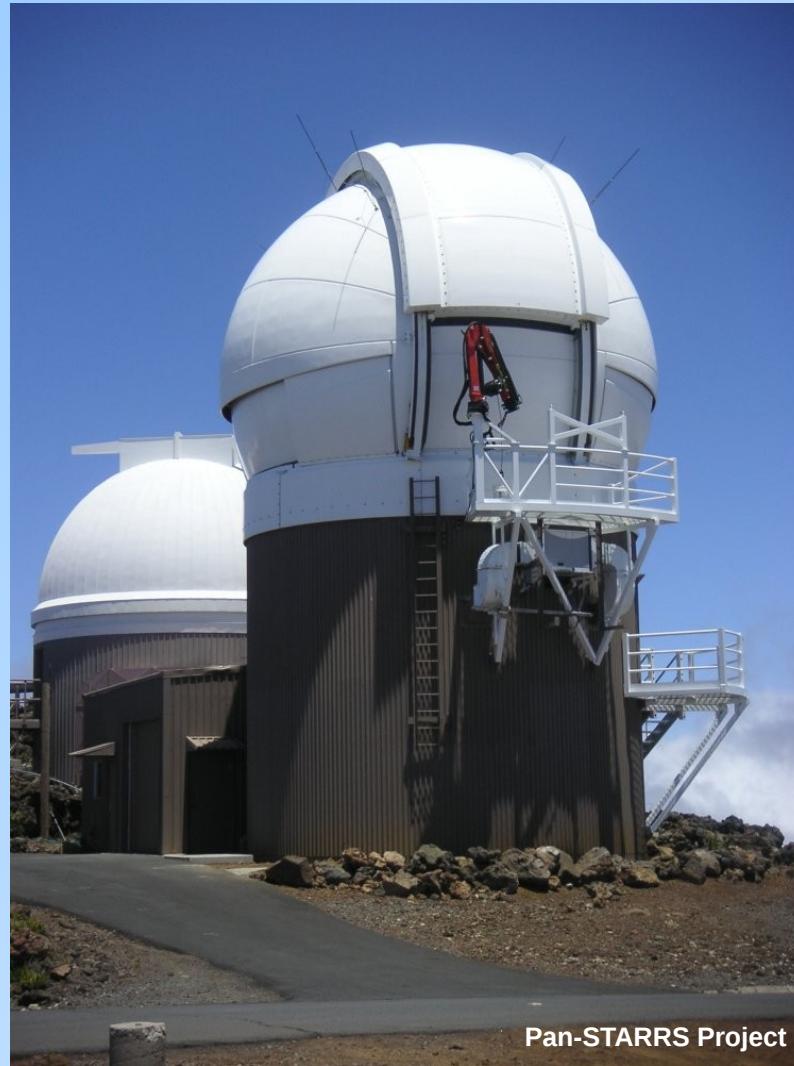


Pan-STARRS 1 Photometry and Astrometry Calibration Status

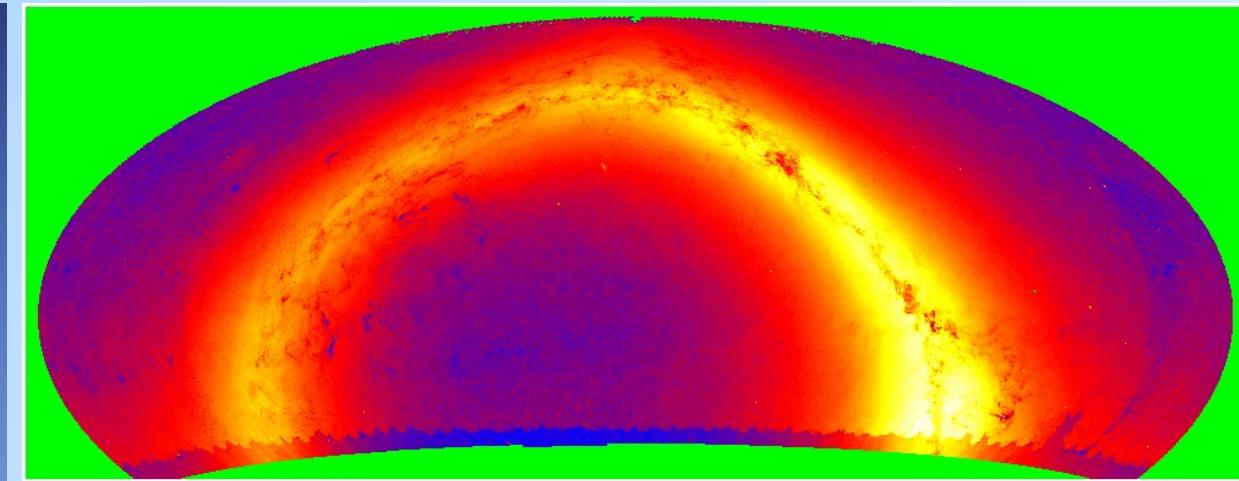
Eugene Magnier
Institute for Astronomy
University of Hawaii



Pan-STARRS 1: a 1.8m survey telescope (1.4Gpix & 7deg² F.O.V.)



PS-1 on Haleakala

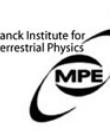


PV1 sky density @ $i = 20.0$

PS1 Survey Mission:

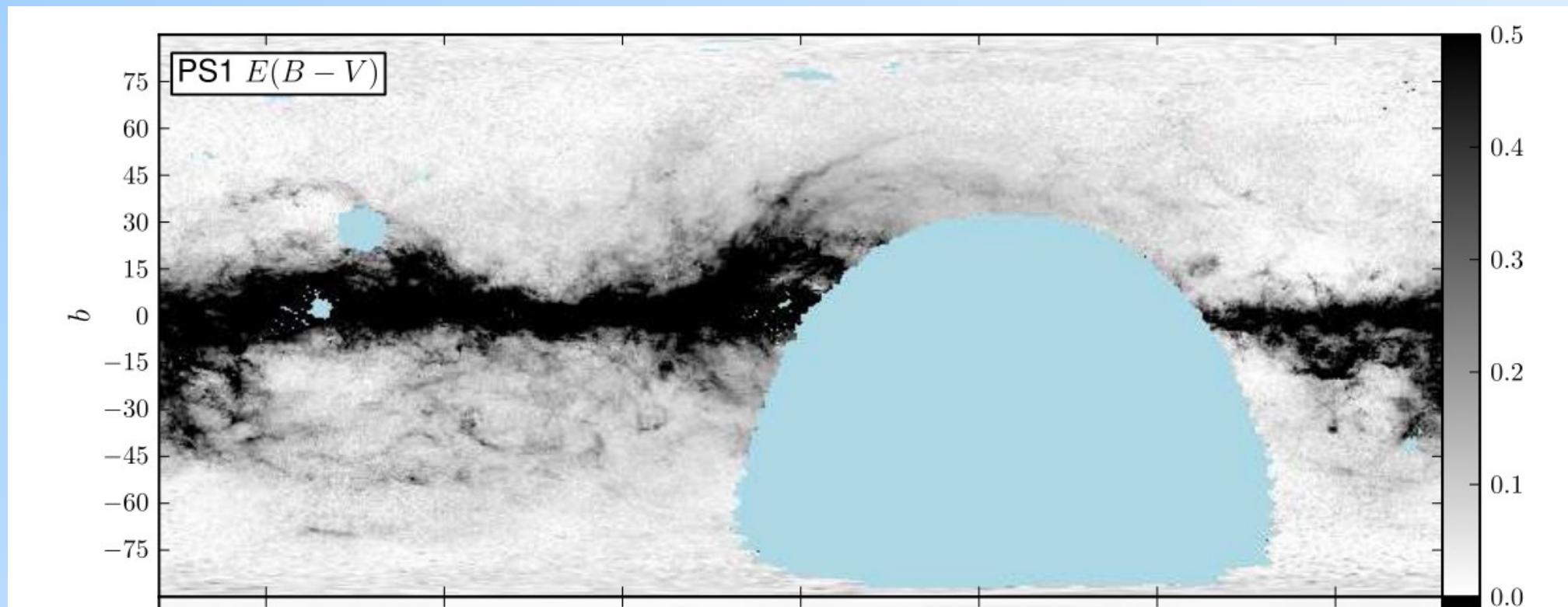
- 5/2010 – 4/2014
- Release to public (MAST): 4/2015 (PV3)
- PV2 stats:
 - 60B chip detections
 - 13B stack detections
 - 7.9B chip 'objects' (grouped detections)
 - 2.8B stack objects
- PV3 processing ongoing (complete ~ Jan 2015)

PS1 consortium members



Precision Photometry Uses: Dust Maps

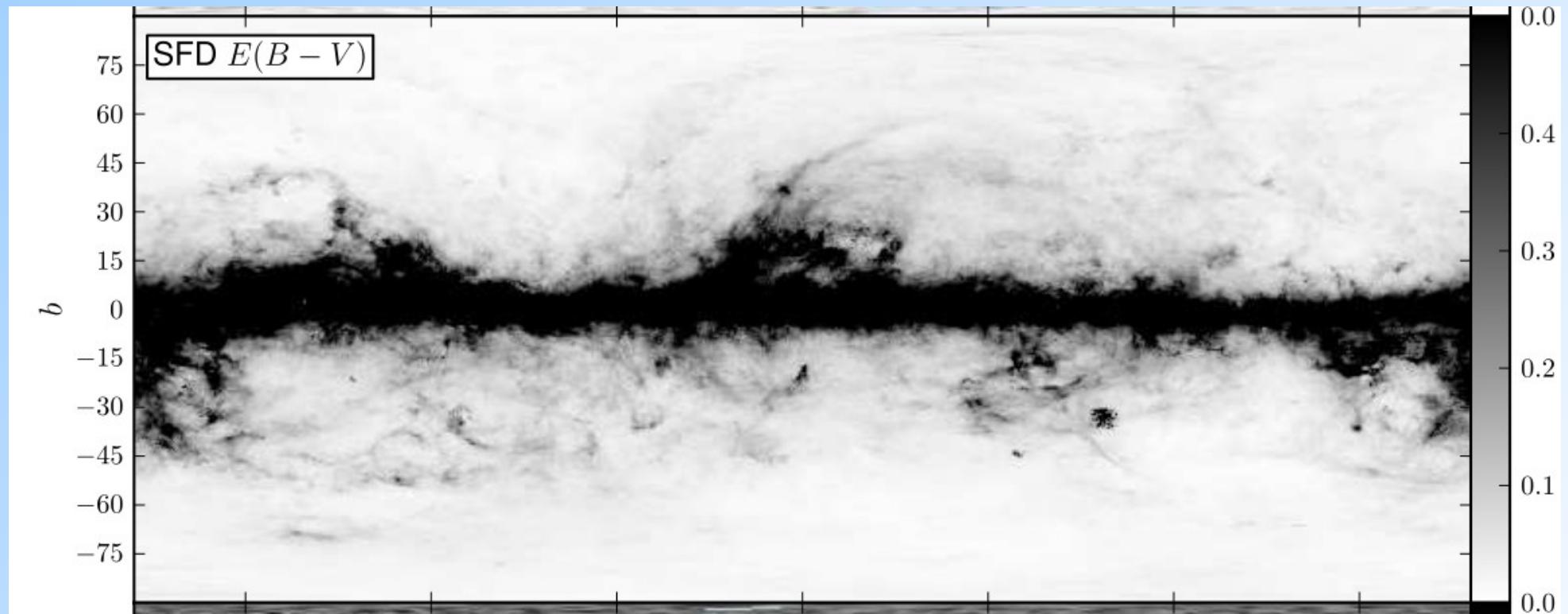
- Schlafly et al (2014) measure $E(B-V)$ trends in lines-of-site



Schlafly et al (2014), ApJ 789, 15

Precision Photometry Uses: Dust Maps

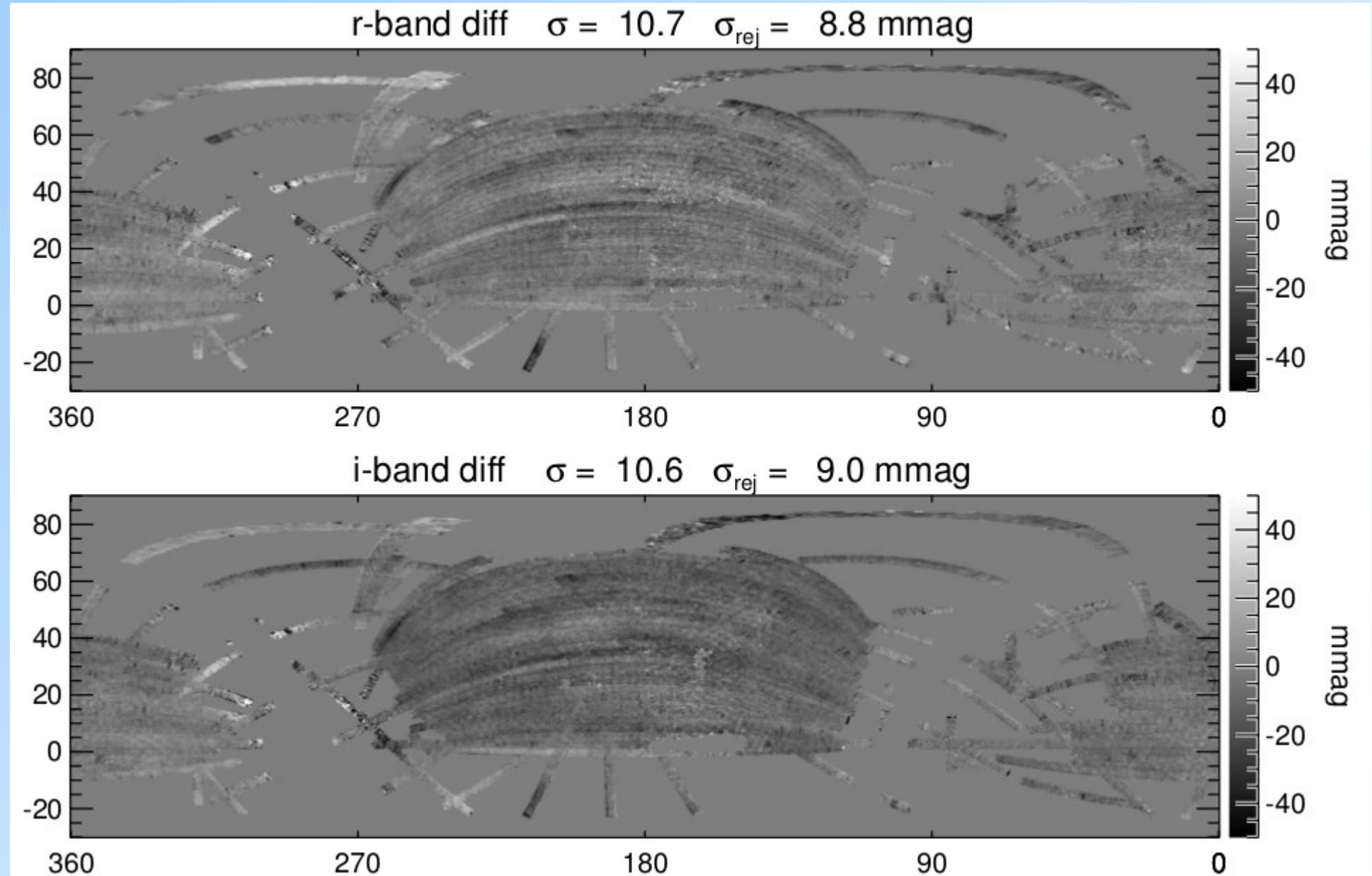
- Schlafly et al (2014) measure $E(B-V)$ trends in lines-of-site
 - Comparison with Schlegel, Finkbeiner, Davis (1998)



Schlafly et al (2014), ApJ 789, 15

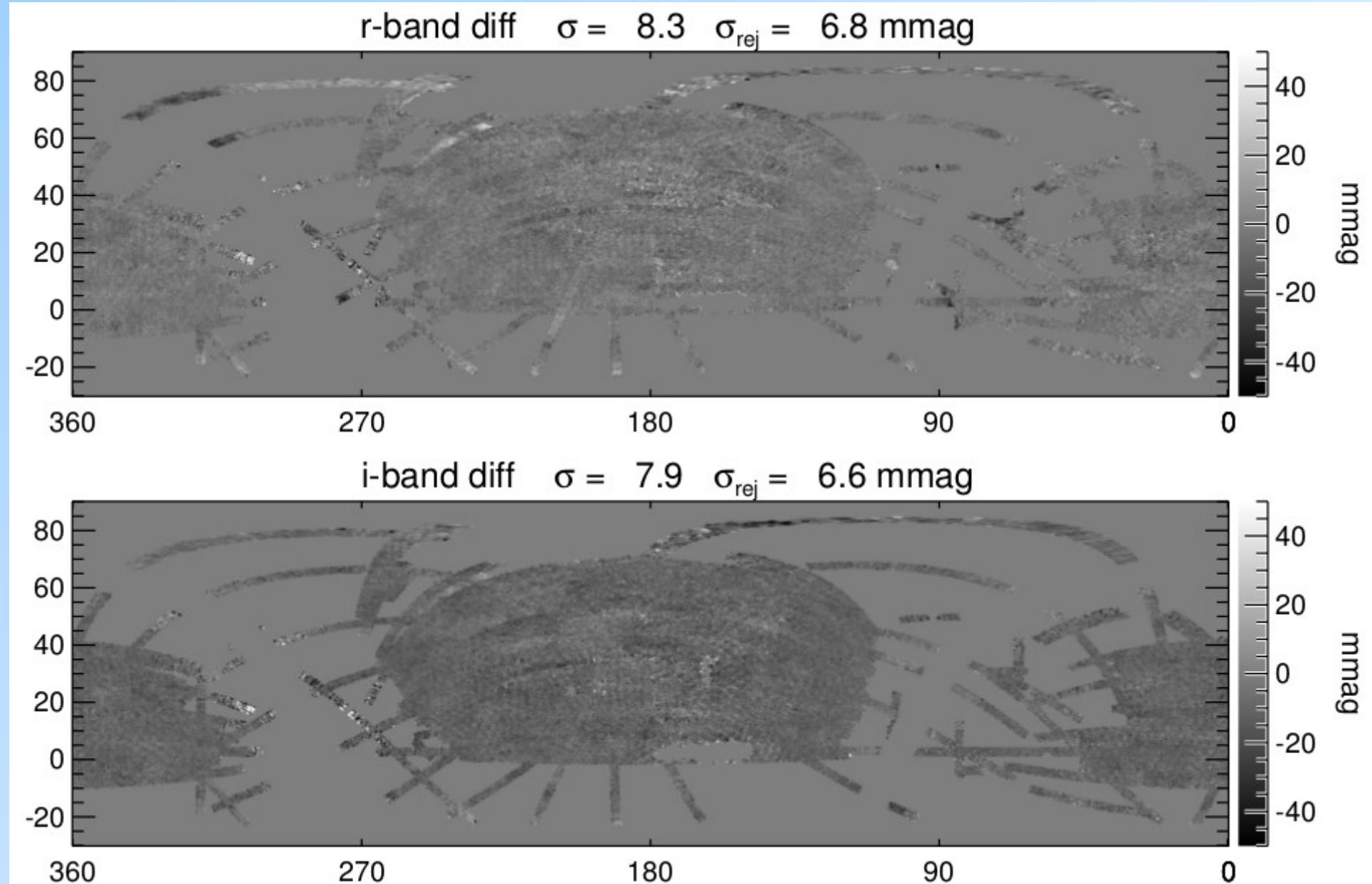
Precision Photometry Uses: SDSS Recalibration

- Finkbeiner et al (2014) uses PS1 to improve the SDSS calibration



Precision Photometry Uses: SDSS Recalibration

- Finkbeiner et al (2014) uses PS1 to improve the SDSS calibration

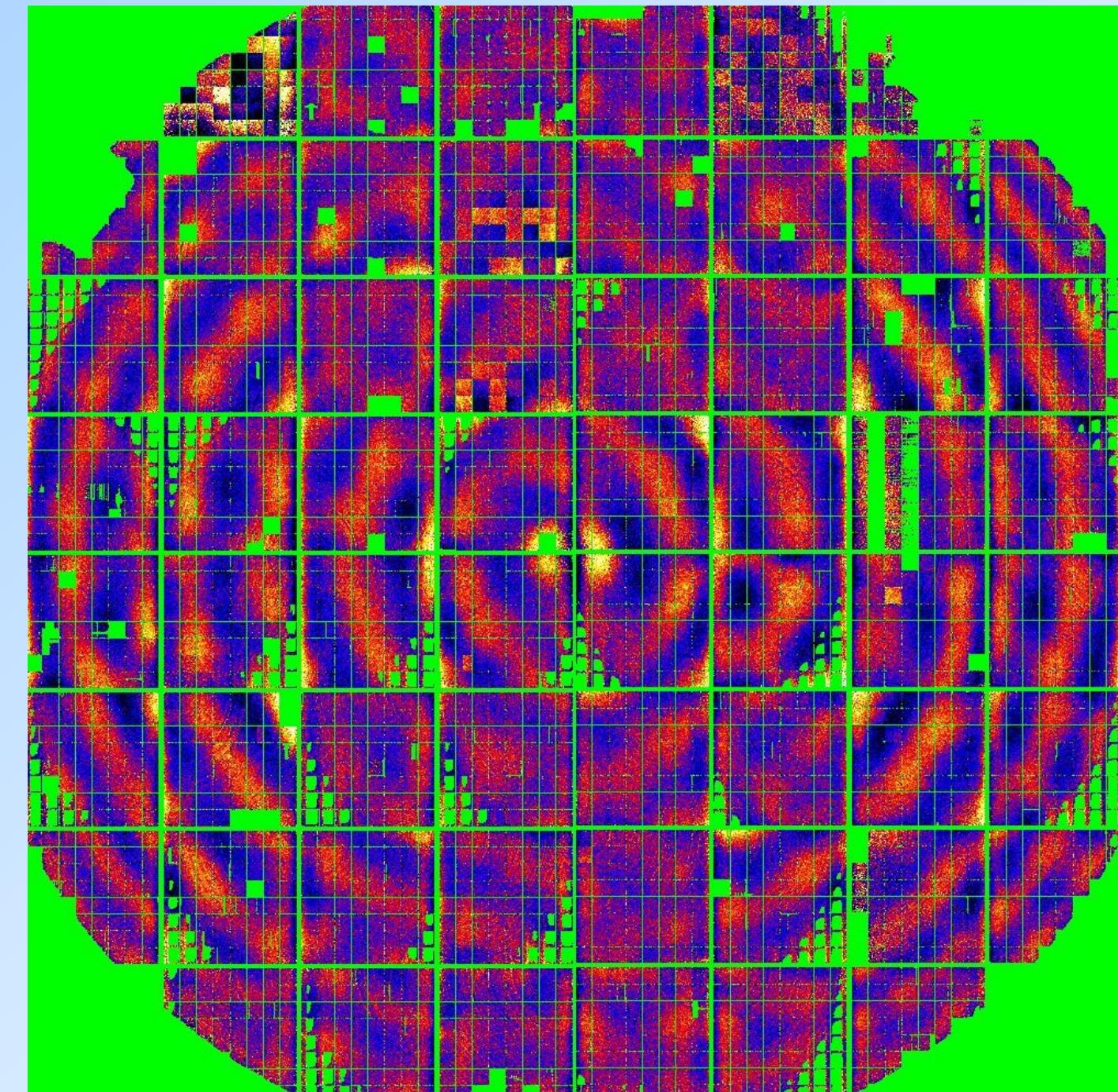


PV1 Calibration Quality

- Photometry
 - 3pi reliability : (*grizy*) = (8, 7, 9, 11, 12) mmag
 - MD reliability : all filters < 6 mmag
 - per-exposure scatter ~ 10 – 15 millimags
- Astrometry:
 - Per-detection : 18 - 20 mas (1D, depends on chip)
 - Parallax limit (1.5yr) : 3-4 mas
 - Proper-motion limit (1.5yr) : ~5 mas/yr
- PV2 & PV3 calibration underway
 - **can we do better?**
 - **probably --> some improvements to pixel processing help**
 - **what about systematics?**
 - **PS1 / GPC1 tree rings**

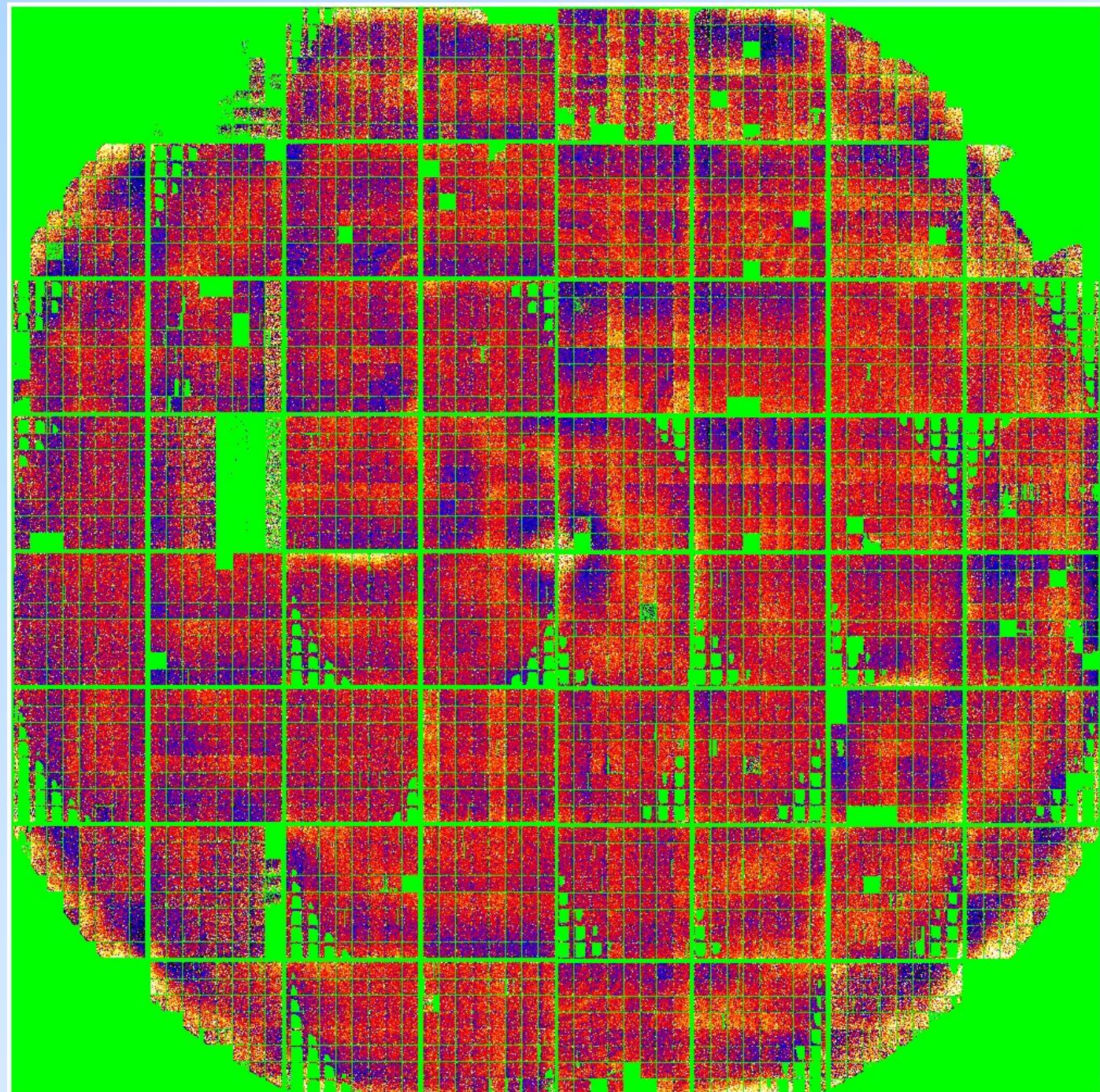
Astrometric Systematics

- mean residuals as a function of camera position
- 20 x 20 pixel bins
- i-band, dX shown
- large-scale structure similar to focal-plane deviations



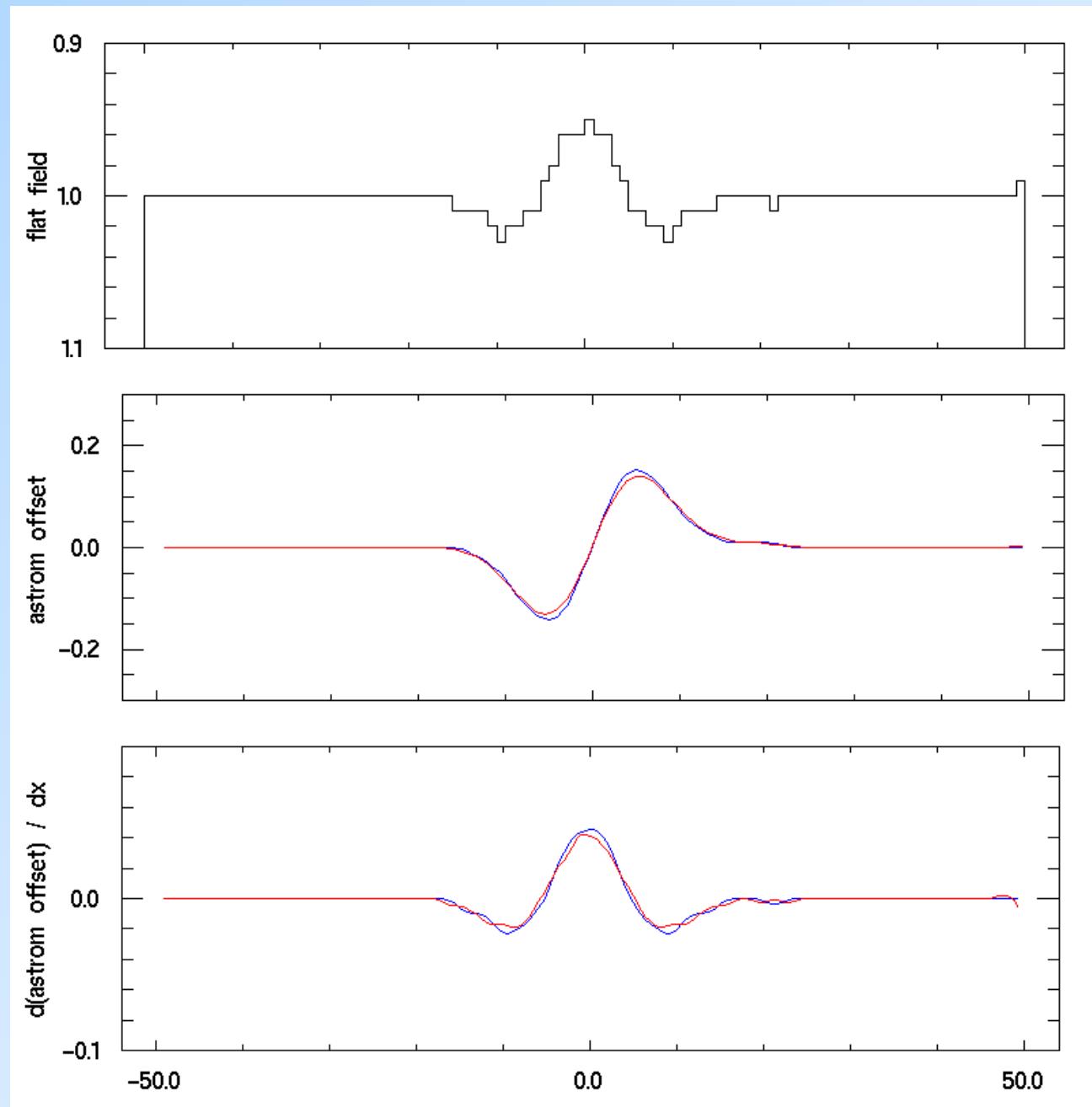
Photometric Systematics : per-exposure residuals

- mean residuals as a function of camera position
- 20 x 20 pixel bins
- i-band shown
- central tent
- residual of 2x2 flats



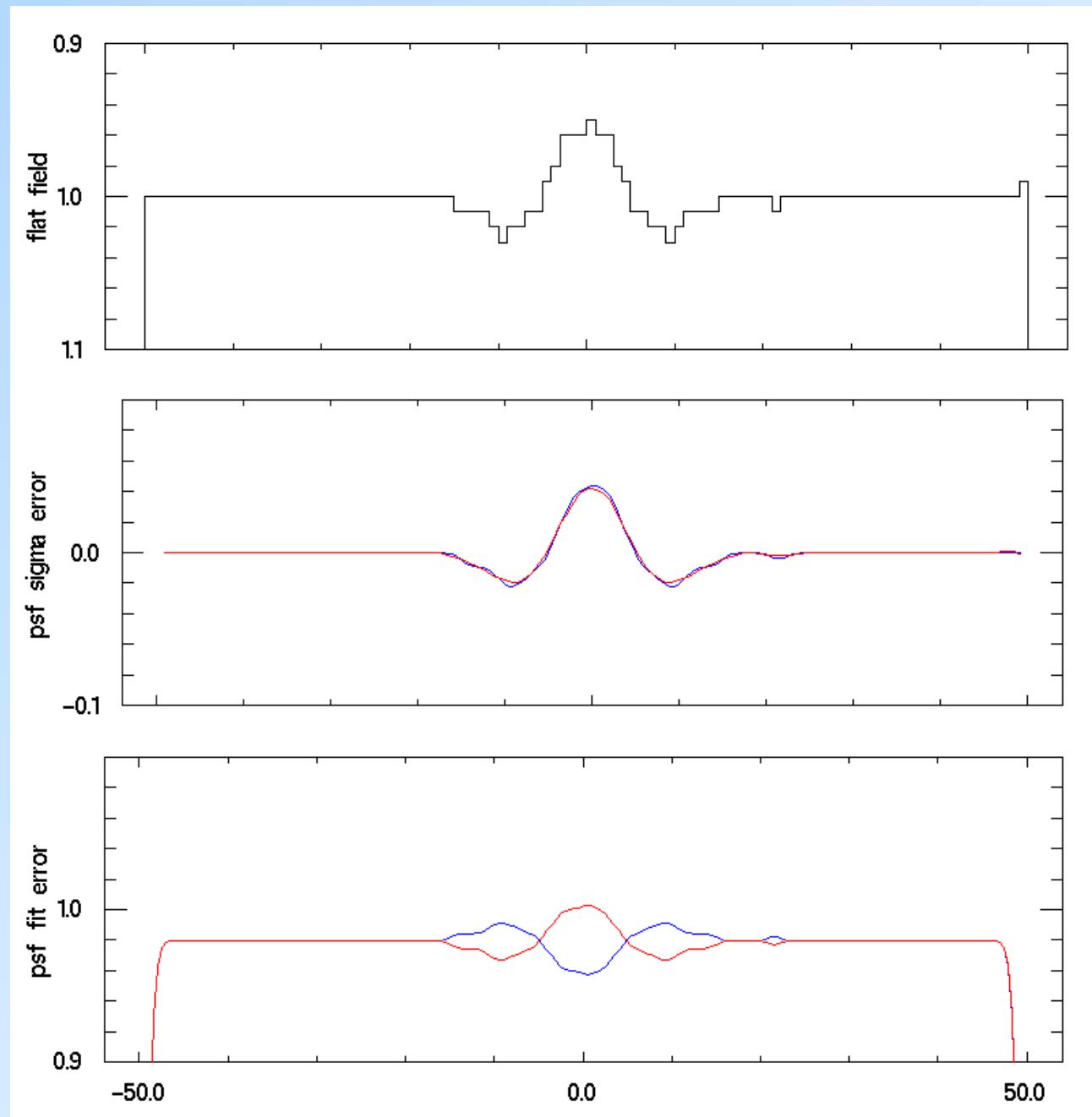
"Classic" Tree Rings : Plate-scale variations

- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)



"Classic" Tree Rings : Plate-scale variations

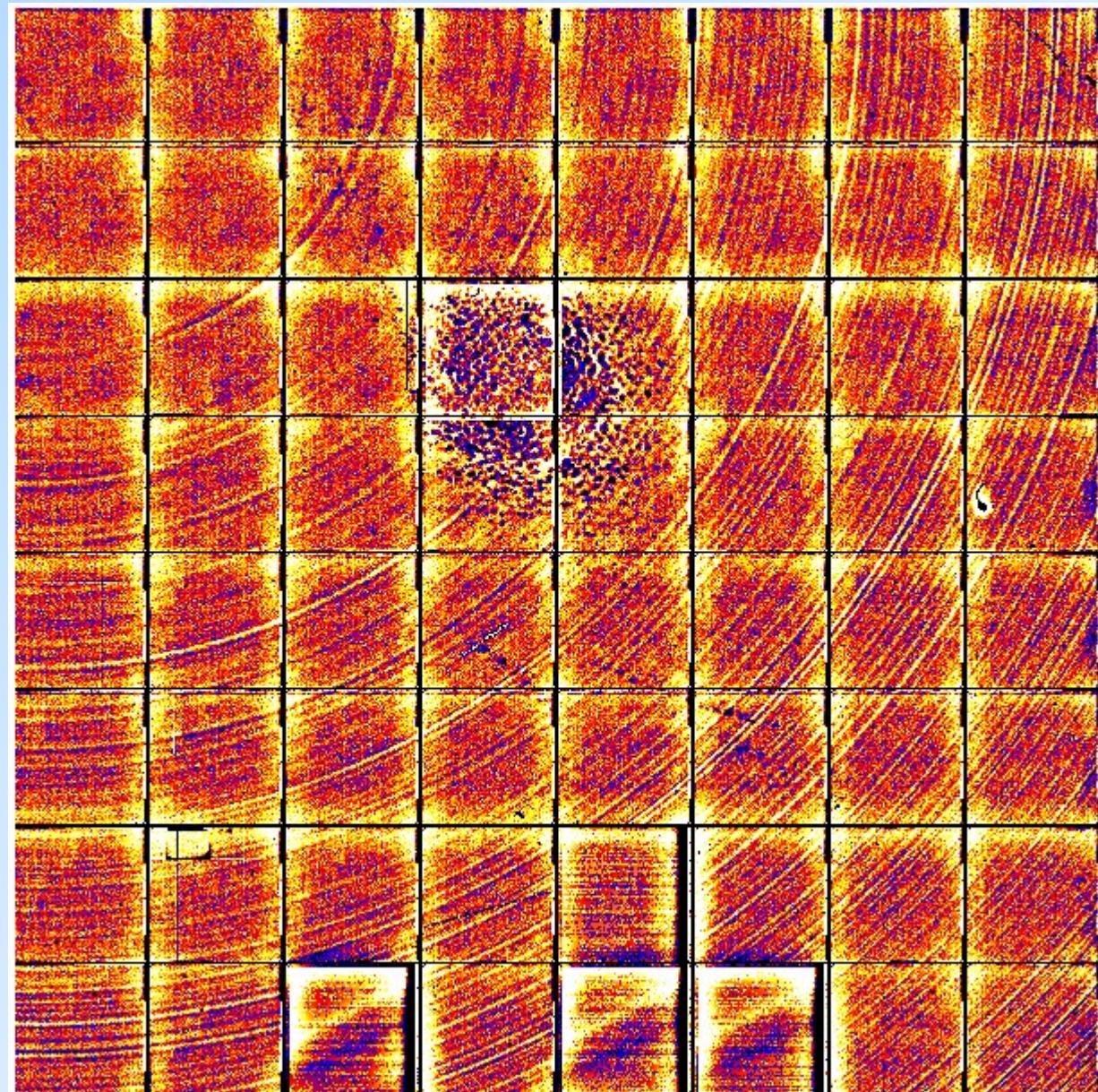
- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)
- plate scale -> psf sigma
- $\Delta_{\text{psf}} \text{ sigma} = \text{grad (astrom)}$
- psfmag err = apmag err / 2



GPC1 Tree Rings : Flat Field (after high-pass filtering)

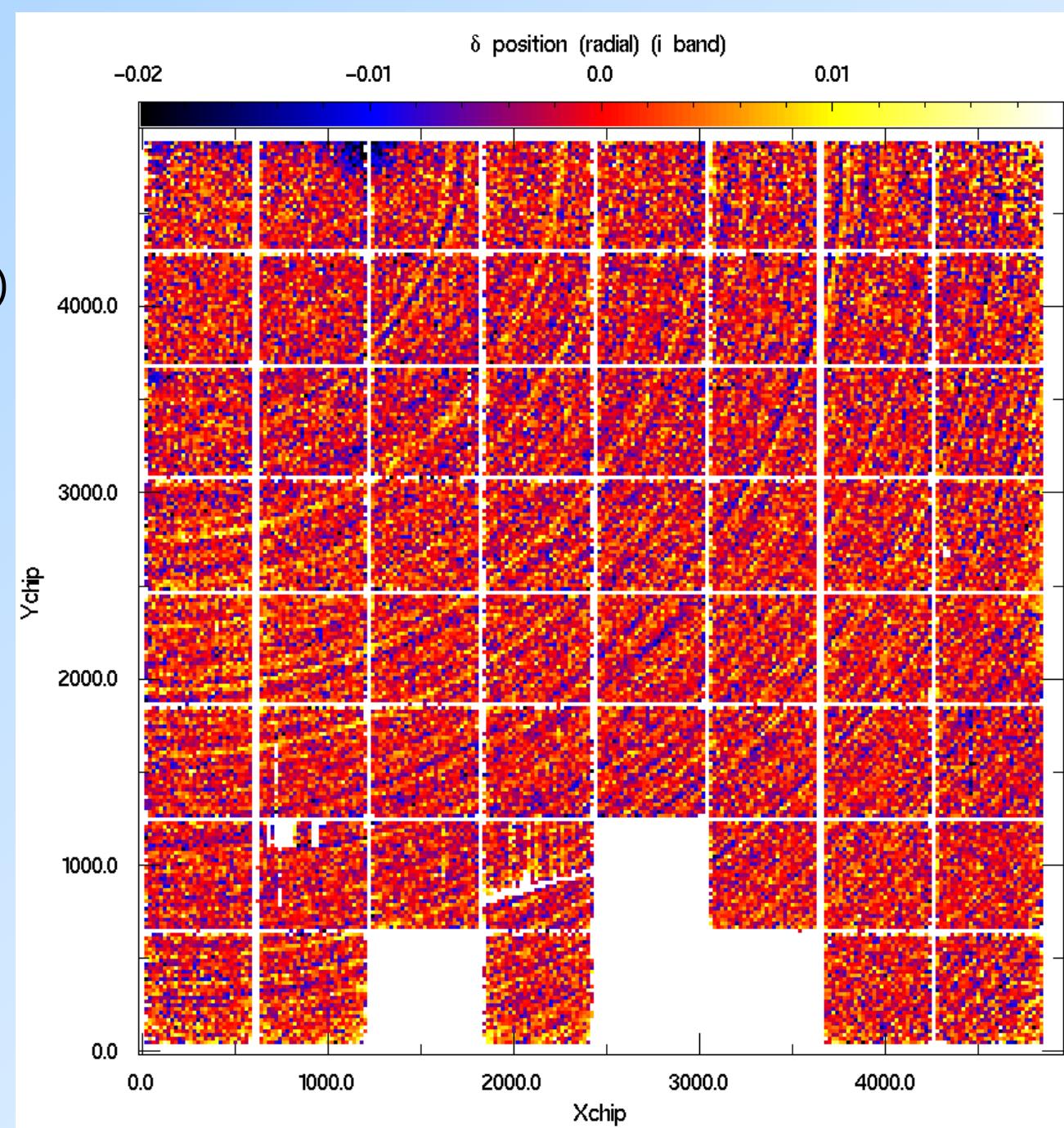
- plate scale -> **flat field**
- plate scale -> astrometry
- flat-field = grad (astrom)
- plate scale -> psf sigma
- Δpsf sigma = grad (astrom)
- psfmag err = apmag err / 2

monochromatic flat @ 630nm



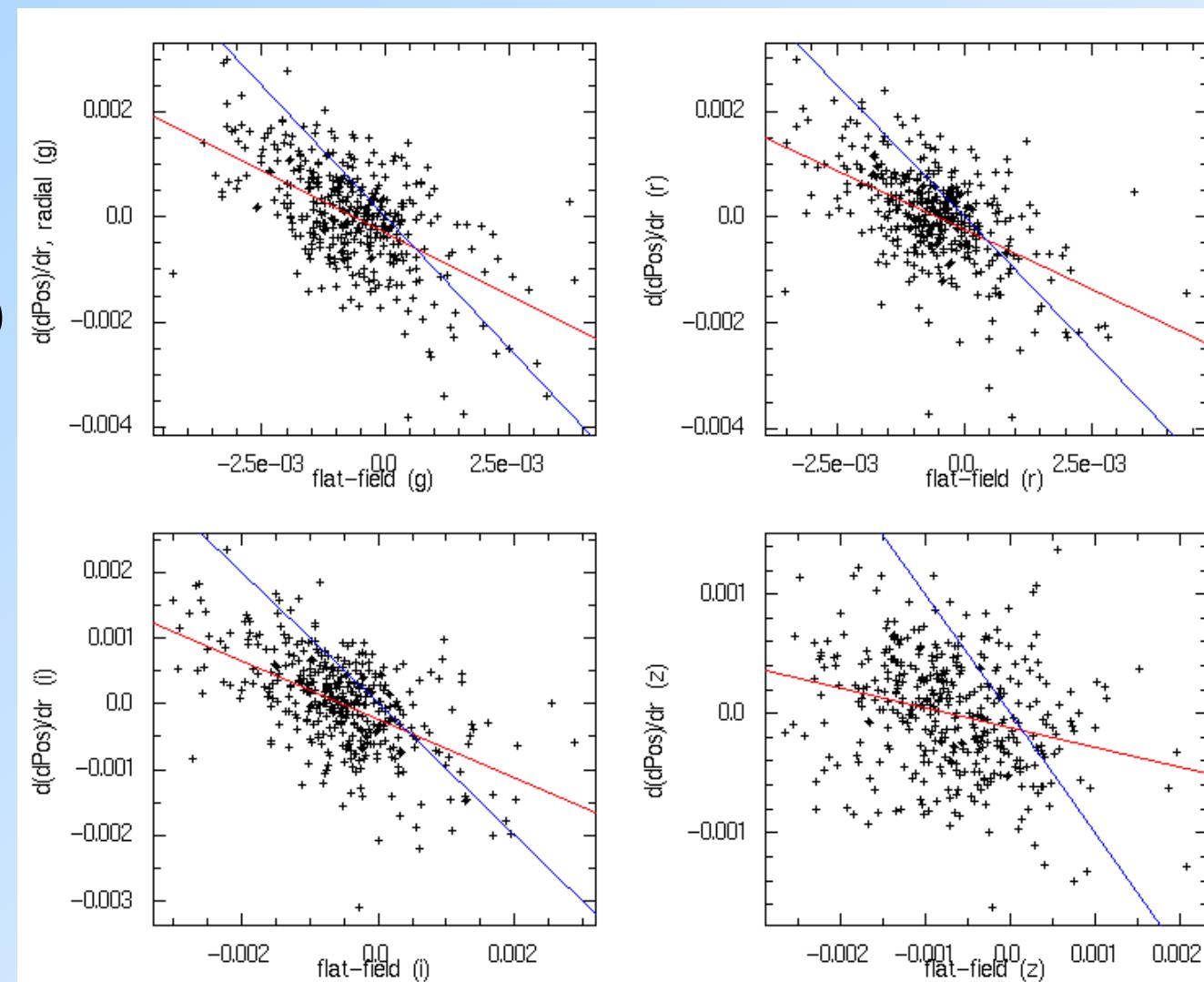
GPC1 Tree Rings : Astrometric Residuals (radial, i-band, XY40)

- plate scale -> **flat field**
- plate scale -> **astrometry**
- flat-field = grad (astrom)
- plate scale -> psf sigma
- Δ psf sigma = grad (astrom)
- psfmag err = apmag err / 2



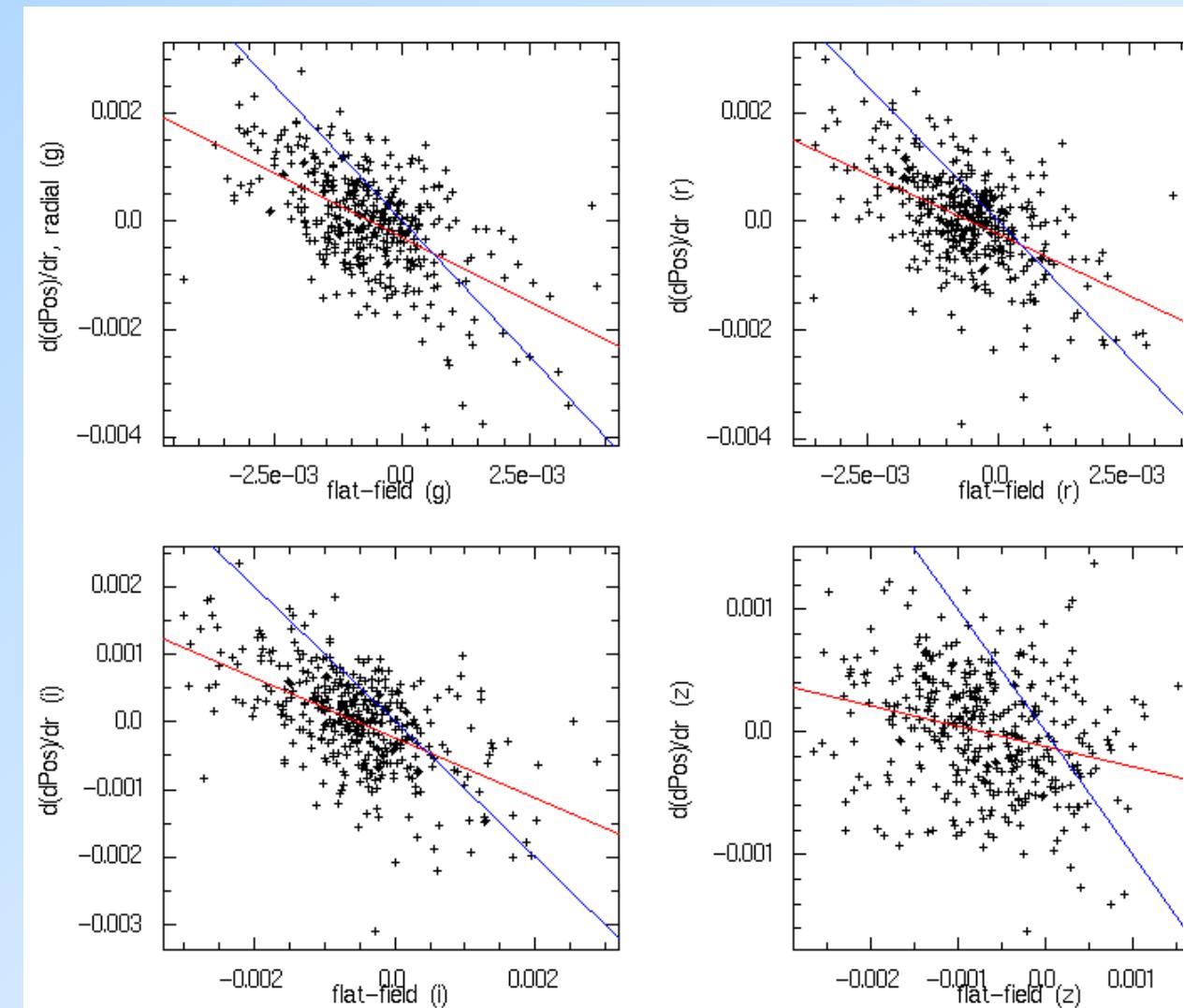
GPC1 Tree Rings : Flat-field = gradient of astrometry errors

- plate scale -> **flat field**
- plate scale -> **astrometry**
- **flat-field = grad (astrom)**
- plate scale -> psf sigma
- Δ psf sigma = grad (astrom)
- psfmag err = apmag err / 2



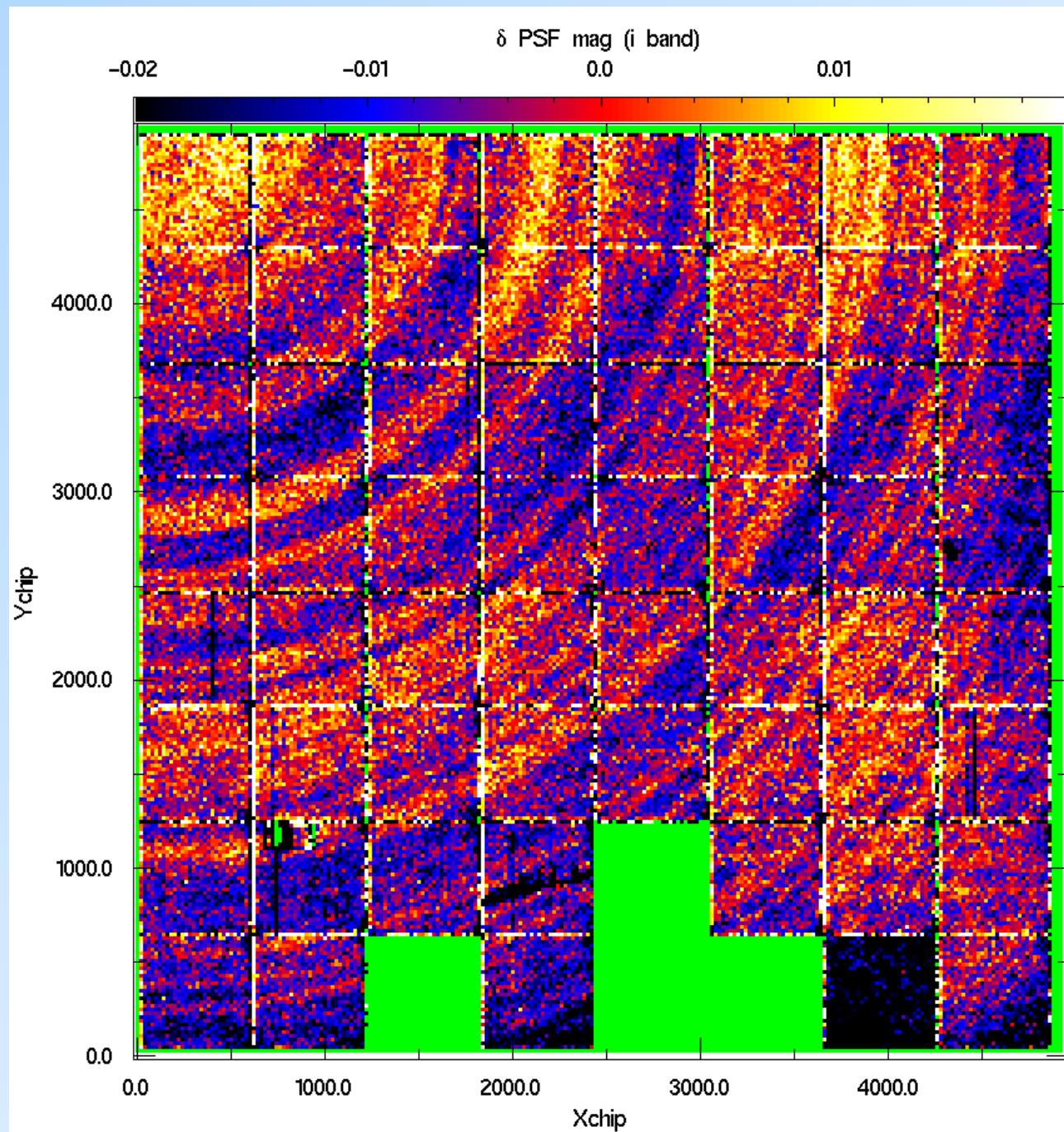
GPC1 Tree Rings : Plate Scale -> Flat-field & astrometry errors

- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)
- plate scale -> psf sigma
- Δ psf sigma = grad (astrom)
- psfmag err = apmag err / 2
- **NOTE amplitude of astrometry error:**
- **4 mas peak-to-peak in GPC1**
- **(vs 15 mas in DEC)**



GPC1 Tree Rings : PSF magnitude residuals

- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)
- plate scale -> psf sigma
- $\Delta_{\text{psf}} \text{ sigma} = \text{grad (astrom)}$
- **psfmag err = apmag err / 2**



GPC1 Tree Rings : Aperture magnitude residuals

- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)
- plate scale -> psf sigma
- Δ psf sigma = grad (astrom)
- **psfmag err = apmag err / 2**

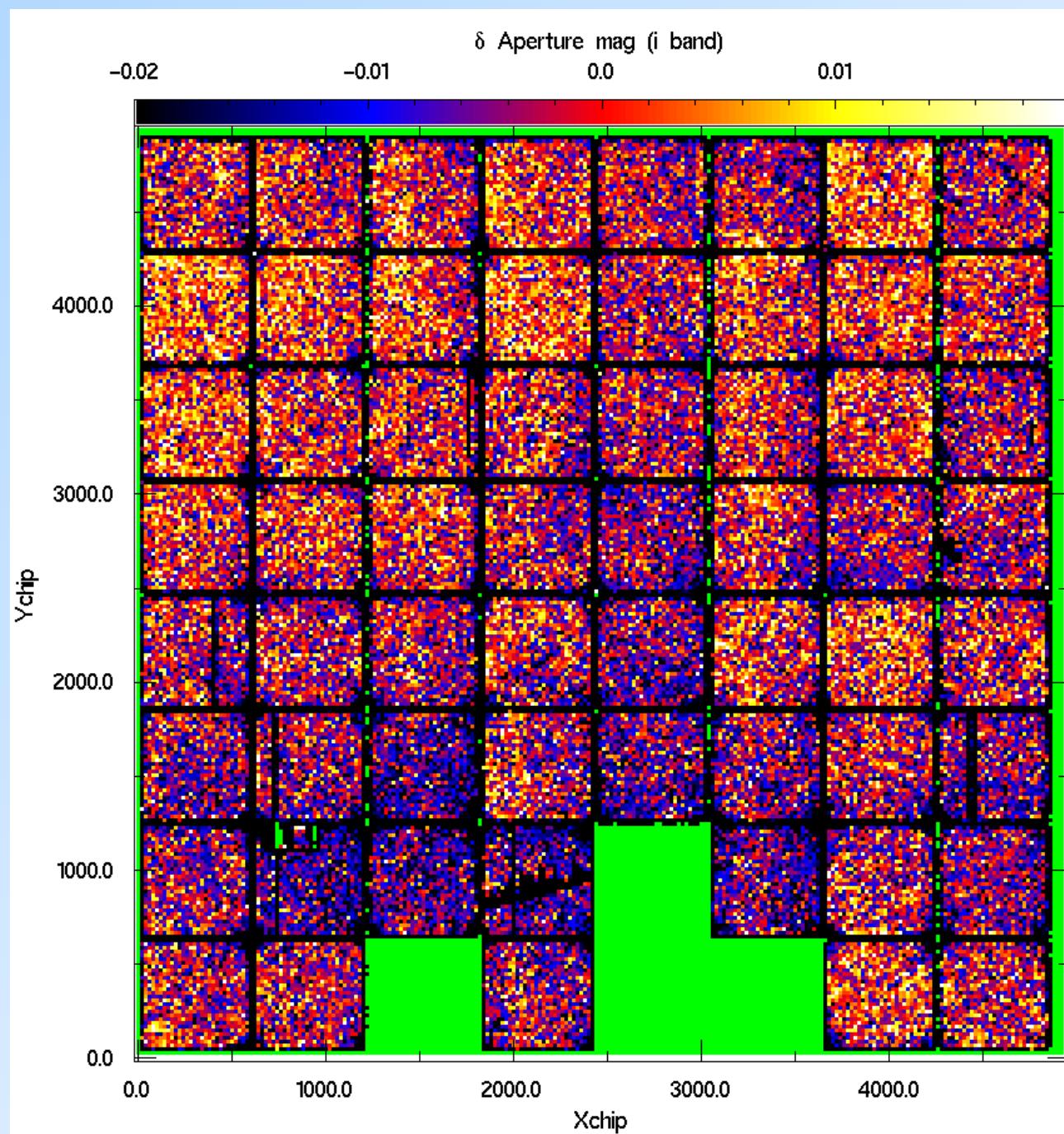


Image Shearing ($\sigma_{\text{major}}^2 - \sigma_{\text{minor}}^2$: i-band, XY40)

- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)
- plate scale -> psf sigma
- Δpsf sigma = grad (astrom)
- psfmag err = apmag err / 2

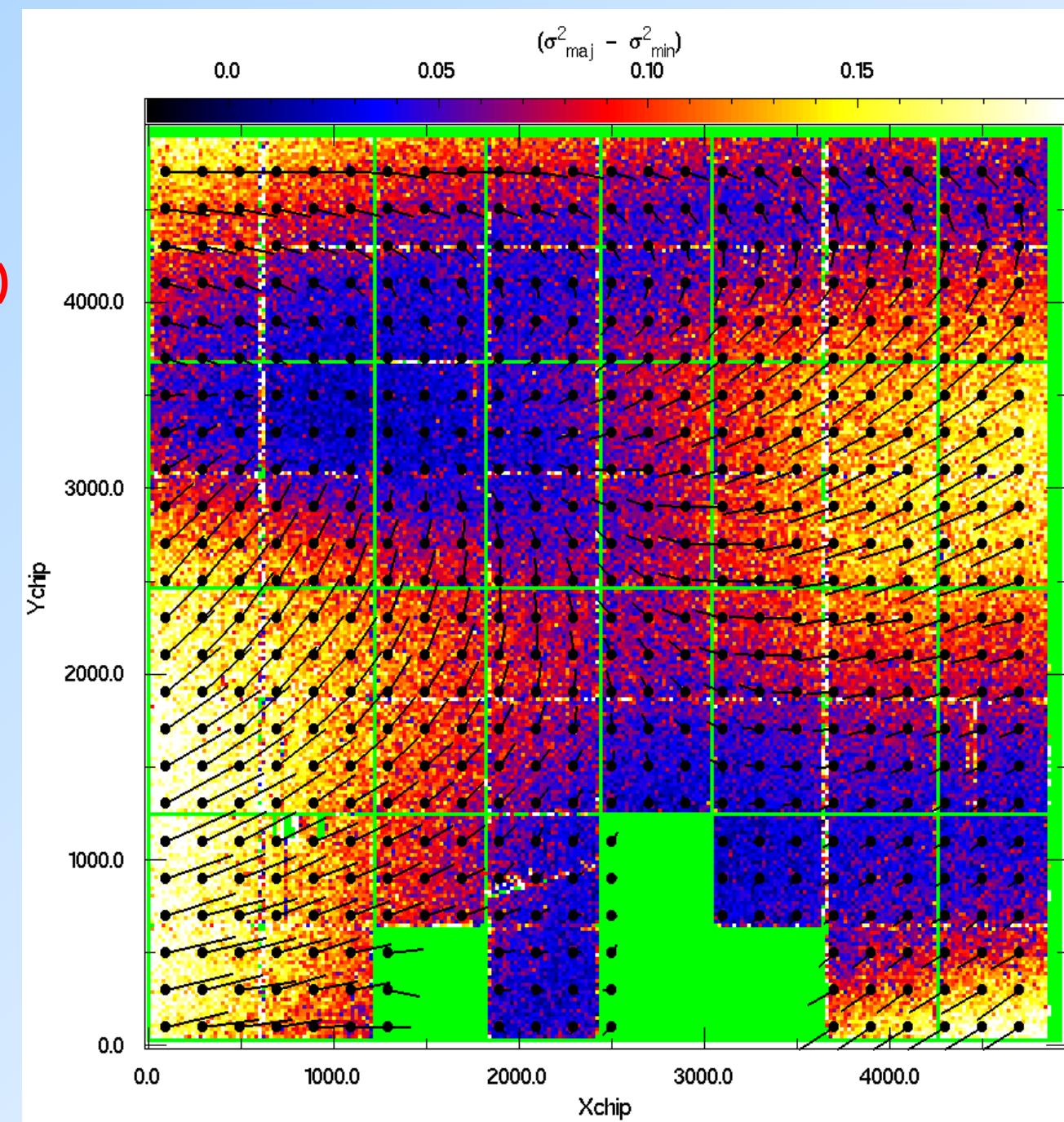
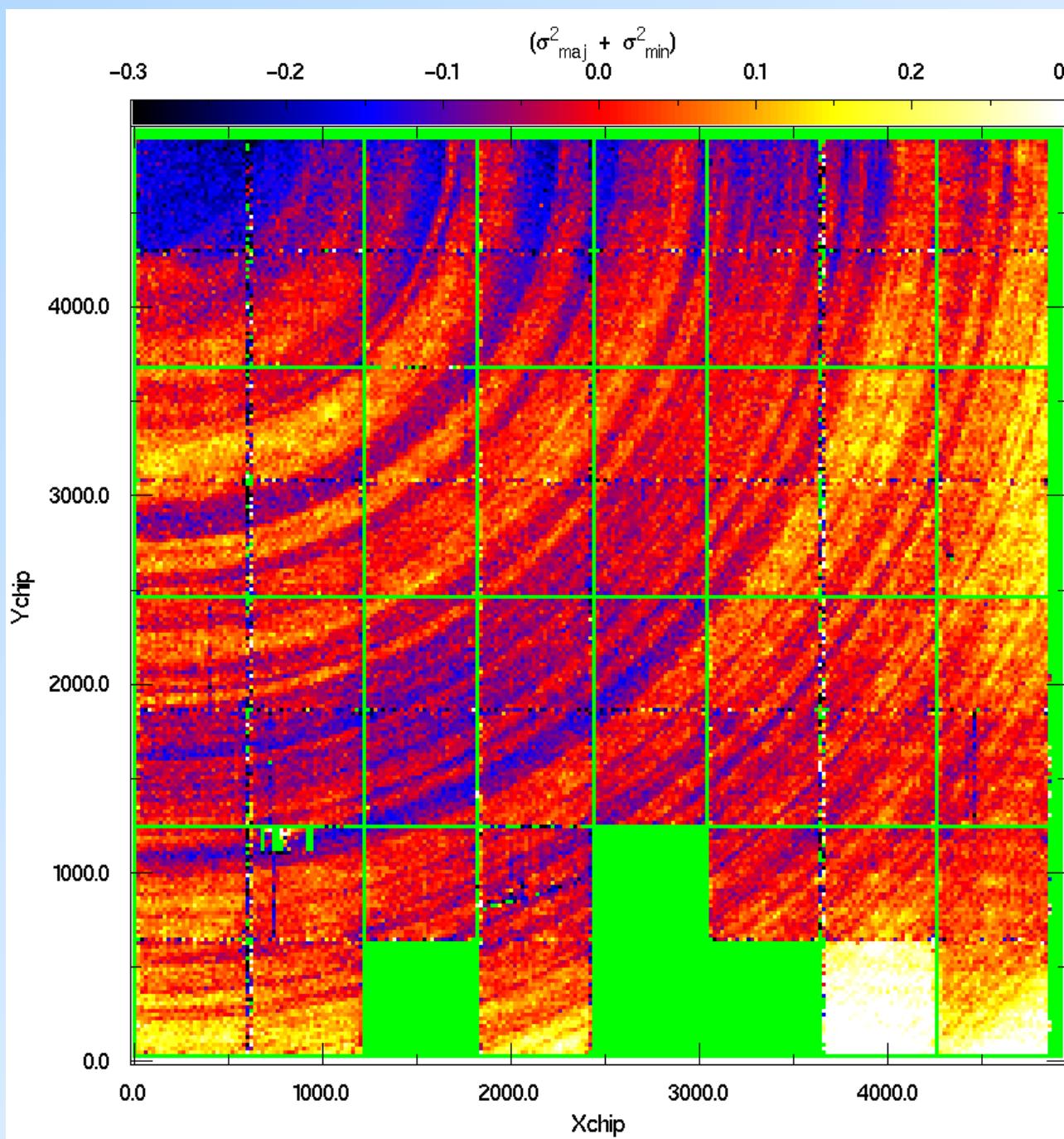


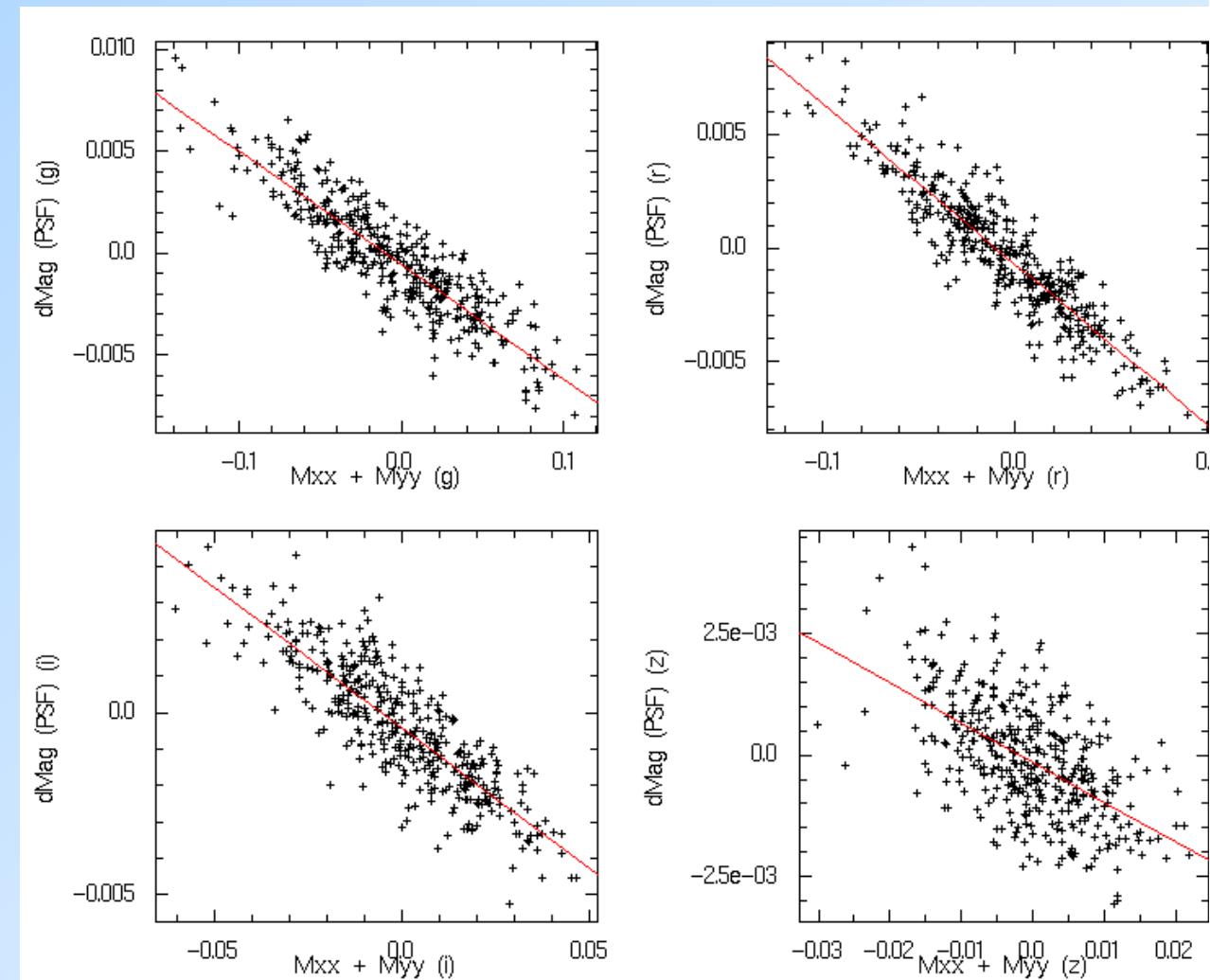
Image Smearing ($\sigma_{\text{major}}^2 + \sigma_{\text{minor}}^2$: i-band, XY40)

- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)
- plate scale -> psf sigma
- Δpsf sigma = grad (astrom)
- psfmag err = apmag err / 2



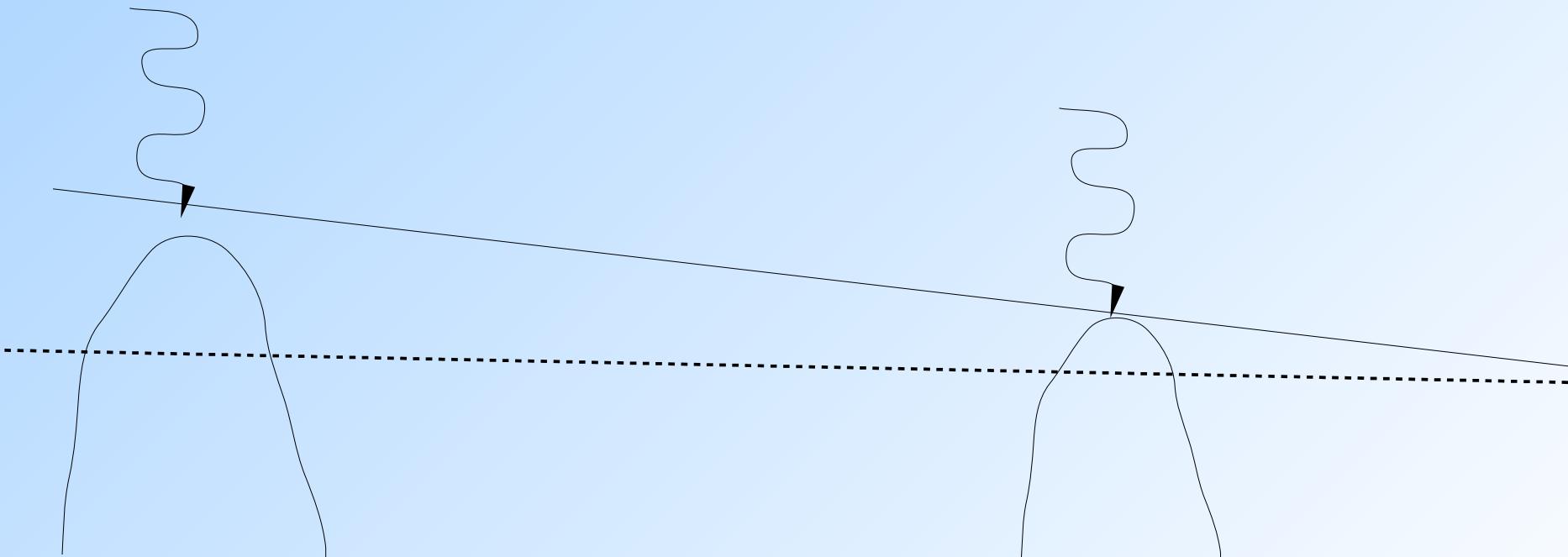
PSF magnitude residuals : correlations between filters

- plate scale -> flat field
- plate scale -> astrometry
- flat-field = grad (astrom)
- ~~plate scale -> psf sigma~~
- ~~Δpsf sigma = grad (astrom)~~
- ~~psfmag err = apmag err / 2~~
- psf smear -> psfmag err
- grad(smear) -> astrom err



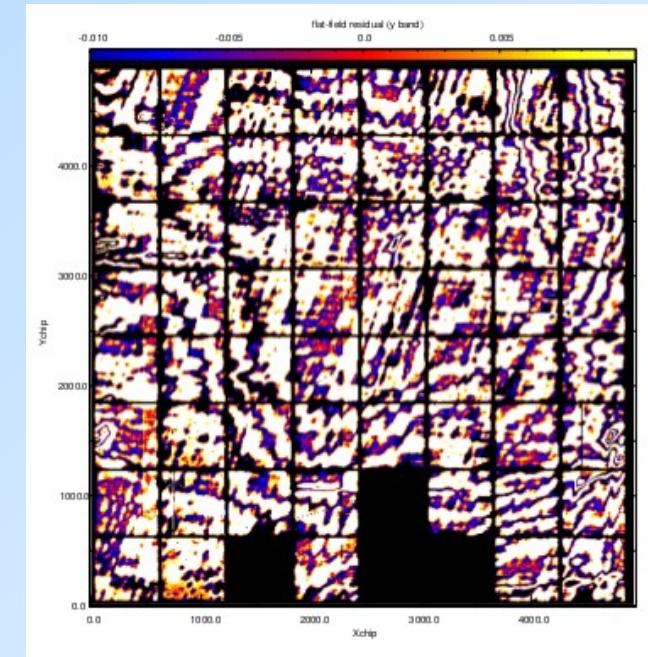
GPC1 Tree Rings : Causes?

- charge diffusion?
 - g-band smearing as an example:
 - $\sigma^2 \sim 0.05 \text{ pixel}^2 \rightarrow \sigma \sim 2.2 \text{ microns}$
 - mean charge diffusion $\sim 3 \text{ microns}$
 - diffusion $\sigma^2 \sim \text{conversion depth [or } E(z)]$
 - if *floor* is 3 microns, range corresponds to 50% variations in depth!
 - for 5 micron variations in depth (1/15), floor is $\sim 8.5 \text{ microns}$



GPC1 Tree Rings : Causes?

- focus variations?
 - $\sigma^2 \sim 0.05 \text{ pixel}^2 \rightarrow \sigma \sim 2.2 \text{ microns}$
 - beam is f/4 $\rightarrow 8.8 \text{ micron variations in depth}$
 - inconsistent with fringe patterns



Conclusions

- GPC1 has a new and exciting type of tree ring effect!
- circularly symmetric smear \rightarrow PSF photometry errors
- related to (weaker) plate-scale variations (traditional tree rings)
- no good model for physical cause
 - charge diffusion implies implausibly large depth variations
 - focus variations implies implausibly large thickness variations
- affects the PSF in quadrature
 - cannot be taken out as multiplicative correction to the mean
 - include in the PSF model or per-exposure measurements
 - may now be the dominant term in PS1 photometry!

Filter	psf photom res vs smear	∂ smear / ∂ radius vs ΔR_{off}	$\partial \Delta R_{\text{off}} / \partial$ radius vs Δ_{flat}
g_{P1}	-0.056	-0.066	-0.47
r_{P1}	-0.071	-0.073	-0.45
i_{P1}	-0.077	-0.095	-0.45
z_{P1}	-0.082	-0.078	-0.17